

Tilting Uranus via the migration of an ancient satellite

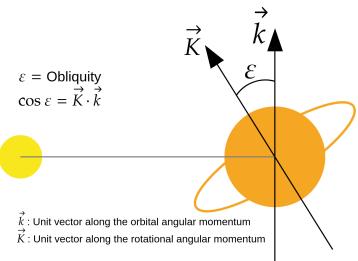
Saillenfest et al. 2022

Jérémy Couturier

November 30<sup>th</sup>, 2023

#### What is the obliquity?







#### Obliquities in the Solar system

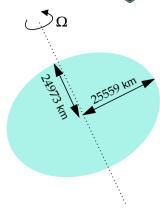
Planet	Obliquity		
Mercury	$0.034^{\circ}$		
Venus	$177.4^{\circ}$		
Earth	$23.44^{\circ}$		
Mars	$25.19^{\circ}$		
Jupiter	$3.13^{\circ}$		
Saturn	$26.73^{\circ}$		
Uranus	$97.77^{\circ}$		
Neptune	$28.32^{\circ}$		

- ▶ Laskar & Robutel (1993): The inner planets' obliquities were chaotic at some point  $\rightarrow$  Not primordial.
- ► Planets are expected to form with near-zero obliquity.
- ► The outer planets' obliquities should be primordial but only Jupiter's is
- ► Why is Uranus lying on its side ?

#### **Equatorial bulge**







- ightharpoonup Uranus is slightly flattened due to its rotation ightarrow equatorial bulge
- ► The pull of the Sun and of the satellite on this bulge yields a torque

How does Uranus react to this torque?

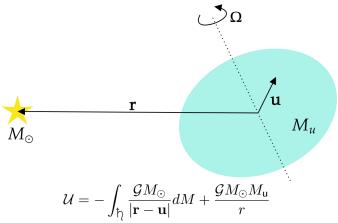


#### Colombo model (Colombo, 1966)

- ▶ I present a simple Hamiltonian version of the Colombo model.
- ▶ The Hamiltonian is written  $\mathcal{H} = \mathcal{T} + \mathcal{U}$  where
- $ightharpoonup \mathcal{T}$  is the kinetic energy of rotation of Uranus on itself
- $ightharpoonup \mathcal{U}$  is the potential energy associated with the interaction between the bulge and the Sun.
- ightharpoonup To only retain the long-term evolution of the rotation,  $\mathcal U$  is averaged over one orbital period of Uranus.

# What about the potential energy ?



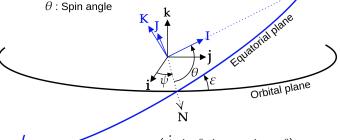


$$\langle \mathcal{U} \rangle := \frac{1}{T} \int_0^T \mathcal{U} dt = -\frac{\mathcal{G} M_{\odot} \left( C - A \right)}{4 a_{\odot}^3 \left( 1 - e_{\odot}^2 \right)^{3/2}} \left( 3 \cos^2 \varepsilon - 1 \right)$$

 ${\cal C}$  and  ${\cal A}$  are the polar and equatorial moment of inertia, respectively

## What about the kinetic energy?





$$\mathbf{\Omega} = \dot{\psi}\mathbf{k} + \dot{\varepsilon}\mathbf{N} + \dot{\theta}\mathbf{K} = \begin{pmatrix} \dot{\psi}\sin\theta\sin\varepsilon + \dot{\varepsilon}\cos\theta\\ \dot{\psi}\cos\theta\sin\varepsilon - \dot{\varepsilon}\sin\theta\\ \dot{\psi}\cos\varepsilon + \dot{\theta} \end{pmatrix}_{\mathbf{I},\mathbf{i}}$$

$$\mathcal{T} = rac{1}{2}\,{}^t \mathbf{\Omega} \mathcal{I} \mathbf{\Omega} \qquad \mathcal{I} = egin{pmatrix} A & 0 & 0 \ 0 & A & 0 \ 0 & 0 & C \end{pmatrix}$$

#### Generalized coordinates and momenta



$$\mathbf{q} = \begin{pmatrix} \psi \\ \varepsilon \\ \theta \end{pmatrix} \qquad \mathbf{p} := \frac{\partial \mathcal{T}}{\partial \dot{\mathbf{q}}}$$

#### Hamilton equations

$$\dot{\mathbf{q}} = \frac{\partial \mathcal{H}}{\partial \mathbf{p}} \qquad \dot{\mathbf{p}} = -\frac{\partial \mathcal{H}}{\partial \mathbf{q}}$$

#### Secular variations of the rotation of Saturn

 $\dot{\theta} = \mathsf{Cste} \ \to \ \mathsf{Length} \ \mathsf{of} \ \mathsf{day} = \mathsf{Cste}$ 

$$\dot{\varepsilon}=0~\rightarrow~{\rm No}~{\rm change}~{\rm in}~{\rm obliquity}$$

$$\dot{\psi} = -\frac{3}{2}\lambda^{-1}\frac{n^2}{\dot{\theta}}\left(1 - e_{\odot}^2\right)^{-3/2}J_2\cos\varepsilon$$

where n is the mean motion of Uranus,  $e_{\odot}$  the eccentricity of its orbit,  $J_2 = (C-A)/M_{\rm u}R_{\rm u}^2$  its second zonal harmonic and  $\lambda = C/M_{\rm u}R_{\rm u}^2$  is the normalized polar moment of inertia of Uranus.



#### Precession of the equinox

The pull of the Sun on the equatorial bulge yields a wobble of the axis of rotation with frequency

$$\dot{\psi} = -p\cos\varepsilon = -\frac{3}{2}\frac{n^2}{\dot{\theta}}\left(1 - e_{\odot}^2\right)^{-3/2}\frac{J_2}{\lambda}\cos\varepsilon$$

For Uranus, the corresponding period is

$$T=rac{2\pi}{i \dot{b}}=22.8~\mathrm{Myr}$$

But what about the contribution from potential satellites ?



**Contribution from a satellite** There are two ways satellites can affect the frequency of the wobble

- ► If the satellite is far away and not in the equatorial plane, it pulls on the bulge like the Sun does.
- ▶ If it is a close-in satellite, it is locked to the equatorial plane and artificially contributes to increase the  $J_2$  of the planet. → It gives leverage to the Sun.

#### Correction of the equinox precession with a satellite

$$\dot{\psi} = -p \left( \cos \varepsilon + \eta \frac{a^2}{r_M^2} \frac{\sin(2\varepsilon - 2I_L)}{2\sin \varepsilon} \right)$$

where  $\eta=m_{\rm s}r_M^2/(2M_{\rm u}J_2R_{\rm u}^2)$  is the dimensionless mass parameter,  $r_M\approx 53R_{\rm u}$  and  $I_L$  is a inclination whose value depends on whether the satellite is in the close-in or far away regime.



The Colombo model applied to Uranus does not predict obliquity variations.

What caused Uranus' obliquity to increase from  $\sim 0^{\circ}$  to  $97.77^{\circ}$ ? Previous studies (e.g. Slattery et al. 1992) suggest that giant impacts during the late heavy bombardment tilted Uranus.

#### Why is this unlikely?

- ► Uranus and Neptune have strikingly similar masses, radii and spin rates, but very different obliquities (98° vs 28°).
- ▶ They also have similar atmospheric dynamics and magnetic fields.
- ▶ One does not expect such similarities from random collisions.

We would expect more diversity if Uranus and Neptune's final formation stages were determined by random collisions.

Let us look for a smoother process · · · A resonance of course !



Planetary motion in the solar system  $\rightarrow$  Forcing on Uranus' spin dynamics.

#### Fourier decomposition of Laskar 1990

$$e_{\odot}e^{i\varpi_{\odot}} = \sum_{k \in \mathbb{Z}} E_k e^{i\left(\mu_k t + \theta_k^{(0)}\right)} \qquad \sin(i_{\odot}/2)e^{i\Omega_{\odot}} = \sum_{k \in \mathbb{Z}} S_k e^{i\left(\nu_k t + \phi_k^{(0)}\right)}$$

 $ightharpoonup e_{\odot}$  : Uranus' eccentricity

▶  $\varpi_{\odot}$ : Uranus' longitude of periapsis

 $ightharpoonup i_{\odot}$  : Uranus' inclination on the ecliptic

 $lackbox{}{}$   $\Omega_{\odot}$  : Uranus' longitude of the ascending node

 $\mu_k$  and  $\nu_k$  are the fundamental frequencies of the Solar system.

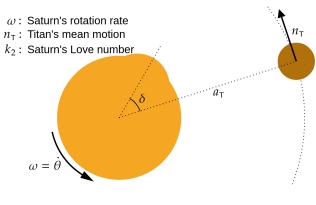


#### Uranus frequencies (Saillenfest et al. 2022 from Laskar 1990)

k	Identification(*)	$v_k$ (" yr <sup>-1</sup> )	$S_k \times 10^9$	$\phi_k^{(0)}$ (°)
1	<i>S</i> <sub>5</sub>	0.00000	13 773 646	107.59
2	$s_7$	-3.00557	8 871 413	320.33
3	$s_8$	-0.69189	563 042	203.96
4	$s_6$	-26.33023	347 710	307.29
5	$-g_5 + g_6 + s_6$	-2.35835	299 979	224.75
6	$-g_5 + g_7 + s_7$	-4.16482	187 859	231.66
7	$g_5 - g_7 + s_7$	-1.84625	182 575	224.56
8	$-g_7 + g_8 + s_8$	-3.11725	59 252	146.97
9	$g_6 - g_7 + s_6$	-1.19906	25 881	313.99
10	$2g_5 - s_7$	11.50319	18 941	101.01
11	$g_5 + g_7 - s_7$	10.34389	11 930	11.68
12	$g_5 - g_6 + s_7$	-26.97744	10 362	225.10
13	$s_1$	-5.61755	10 270	348.70
14	$-g_5 + g_6 + s_7$	20.96631	7346	237.78
15	$g_7 - g_8 + s_7$	-0.58033	5474	197.32

#### Why do satellite migrate?

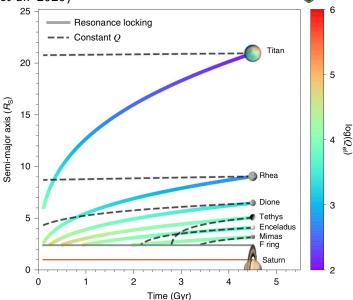




$$\frac{\dot{a}_{\mathsf{T}}}{a_{\mathsf{T}}} = 3\frac{k_2}{Q} \frac{m_{\mathsf{T}}}{M_{\mathsf{s}}} \left(\frac{R_{\mathsf{s}}}{a_{\mathsf{T}}}\right)^5 n_{\mathsf{T}} \qquad Q = \frac{1}{\tan \delta}$$

# Fast tidal migration of Saturnian moons (Lainey et al. 2020)

ROCHESTER ROCHESTER





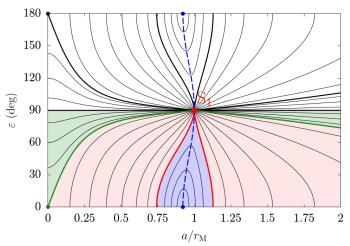
#### Mechanism of tilting

$$\dot{\psi} = -p \left(\cos \varepsilon + \eta \frac{a^2}{r_M^2} \frac{\sin(2\varepsilon - 2I_L)}{2\sin \varepsilon}\right) = \nu_k \approx \mathsf{Cst}$$

- $\blacktriangleright$  If the semi-major axis a of the ancient satellite increases due to tides
- ▶ Then  $\cos \varepsilon$  must decrease to maintain the relation.
- $\blacktriangleright$  The level lines of  $\dot{\psi}$  are followed during the tilting

# Level lines of $\dot{\psi}$





#### Pink region

$$p \le |\nu_k| \le p \frac{\eta}{2}$$

#### Minimal mass for a satellite able to tilt Uranus

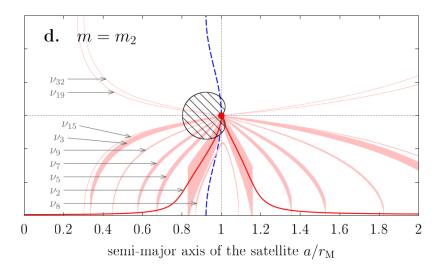


$$|\nu_k| \le p \frac{\eta}{2} = \frac{p}{4} \frac{m_s}{M_{\rm H}} \frac{r_M^2}{J_2 R_{\rm H}^2}$$

k	Identification	$(" \text{yr}^{-1})$	T <sub>lib</sub> (Myr)	$\frac{m_{\min}/M}{(\times 10^{-5})}$	$\frac{m_k/M}{(\times 10^{-5})}$
15	$g_7 - g_8 + s_7$	-0.58033	3021	35	36
3	<i>S</i> <sub>8</sub>	-0.69189	115	41	44
9	$g_6 - g_7 + s_6$	-1.19906	519	71	79
7	$g_5 - g_7 + s_7$	-1.84625	92	110	129
5	$-g_5 + g_6 + s_6$	-2.35835	51	141	172
2	<i>S</i> 7	-3.00557	4	179	233
8	$-g_7 + g_8 + s_8$	-3.11725	115	186	244
6	$-g_5 + g_7 + s_7$	-4.16482	40	248	368

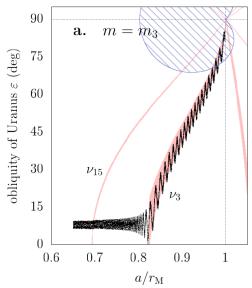
# Resonance $\dot{\psi} = s_7 \ (m_2 = 2.3 \times 10^{-3} M_{\rm u})$





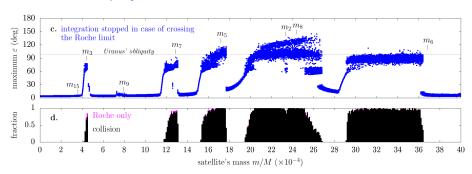
# Example of tilting to a large obliquity







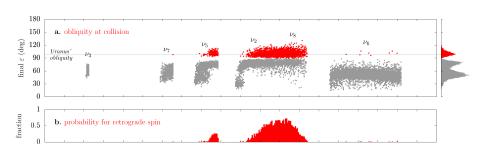
## Maximum obliquity as a function of $m_{\rm s}$



Resonances 2, 5, 6 and 8 look to be potential candidates

# Obliquity at collision as a function of $m_{\rm s}$



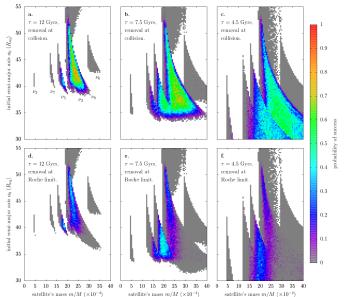


Resonances 2 and 8 not only can reproduce Uranus' obliquity, but that is their most likely outcome !

- $lackbox{ Resonance 2}:\dot{\psi}pprox s_7:$  Forcing due to Uranus' ascending node
- ▶ Resonance 8 :  $\dot{\psi} \approx s_8 + g_8 g_7$  : Forcing due to Neptune's ascending node and Neptune and Uranus' periapsis.

#### Probability of success







#### Conclusion

- ► Provided that enough massive ancient satellite existed, this is a likely mechanism
- ▶ The tilting could also have been due to several satellite

Thank you for your attention